

# The RFP: Its Confinement Status and Future

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## Outline

- Status of RFP confinement
- MST opportunities
  - For general science requiring a PoP facility
  - As a PoP fusion facility

## Key confinement result

*RFP confinement is comparable to tokamak confinement*

*(although RFP confinement improvement is transient)*

# The RFP Confinement Problem

Weak toroidal magnetic field



$$q < 1$$



strong magnetic fluctuations



stochastic field



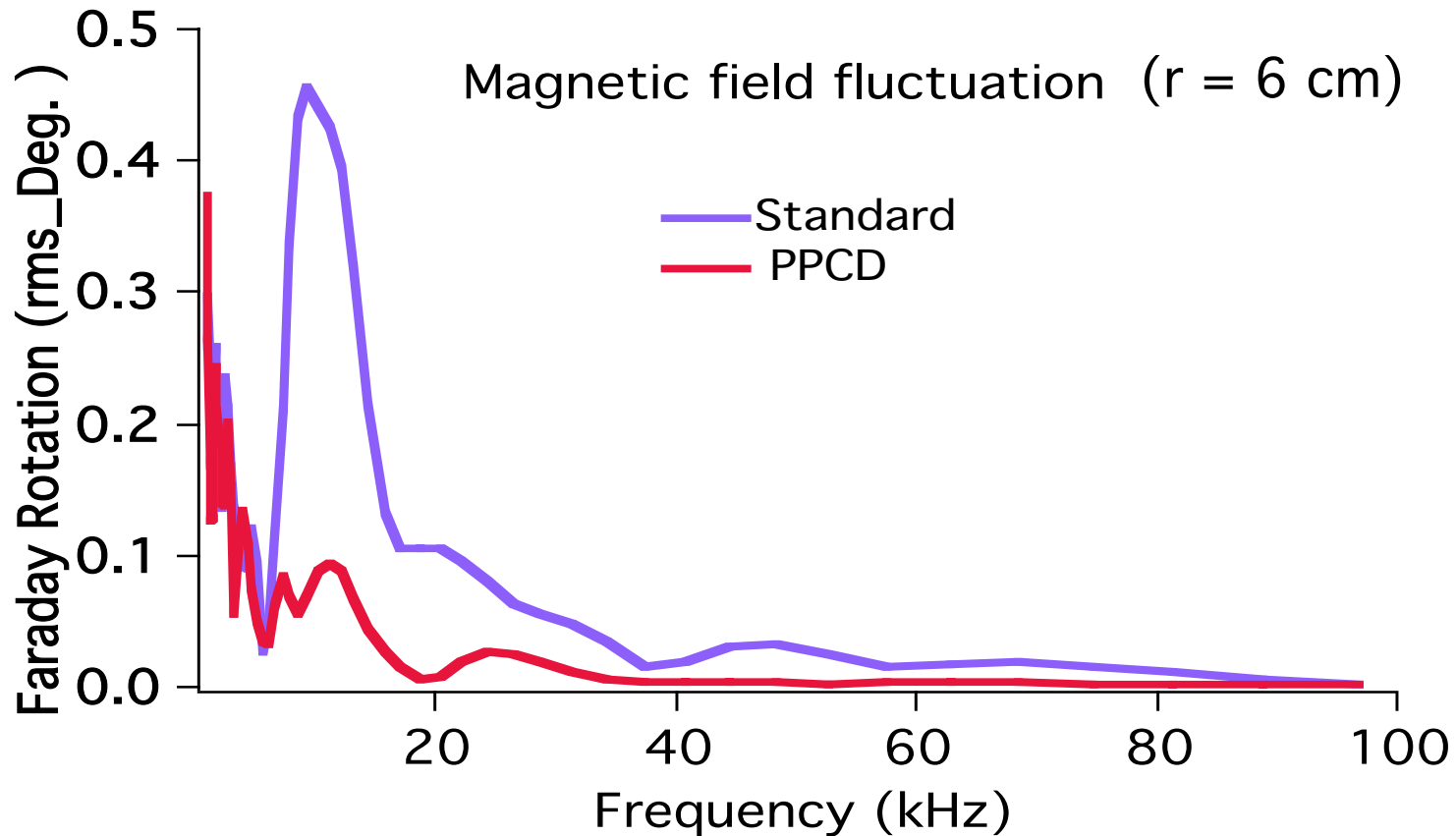
**transport**

## The Solution

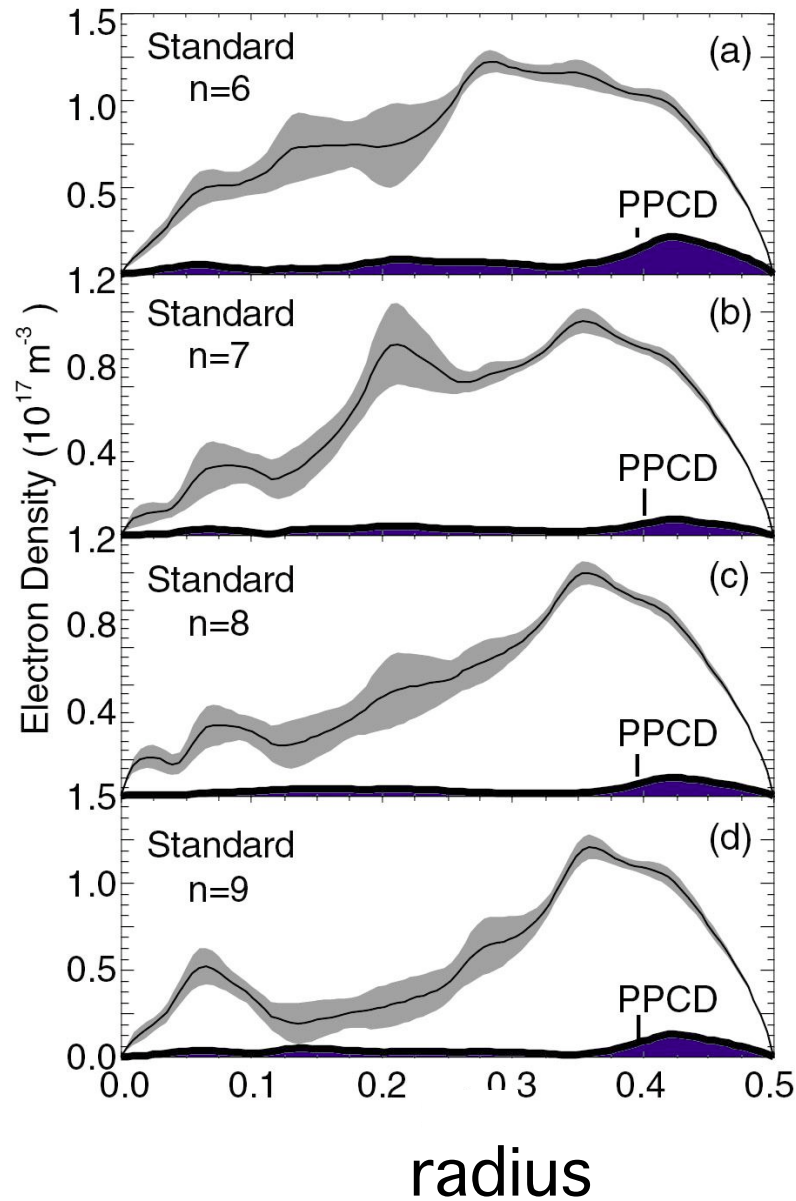
- Control the  $j(r)$  profile to reduce magnetic fluctuations
- First implementation: program Ohmic electric field (transient, crude)

# Core magnetic fluctuations reduced

Measured by Faraday rotation (UCLA)

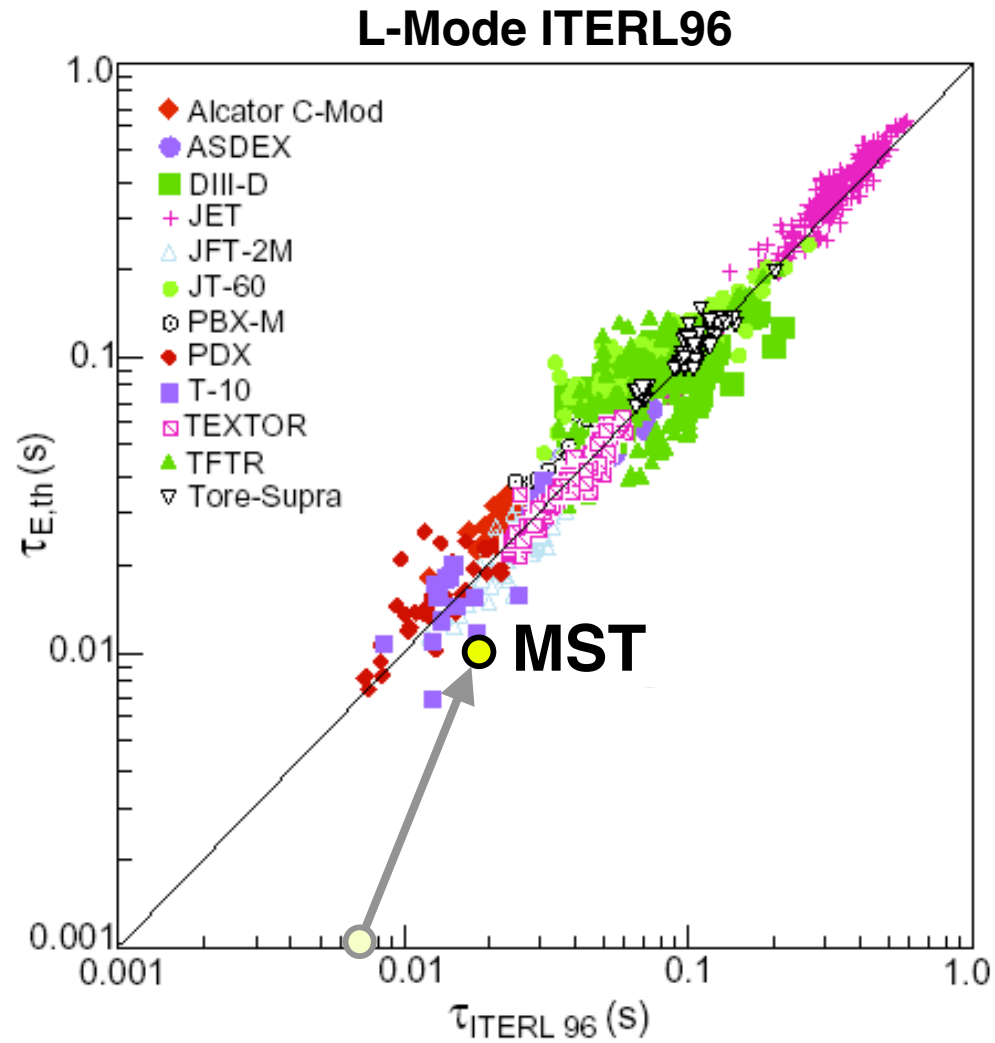


# Core density fluctuations reduced



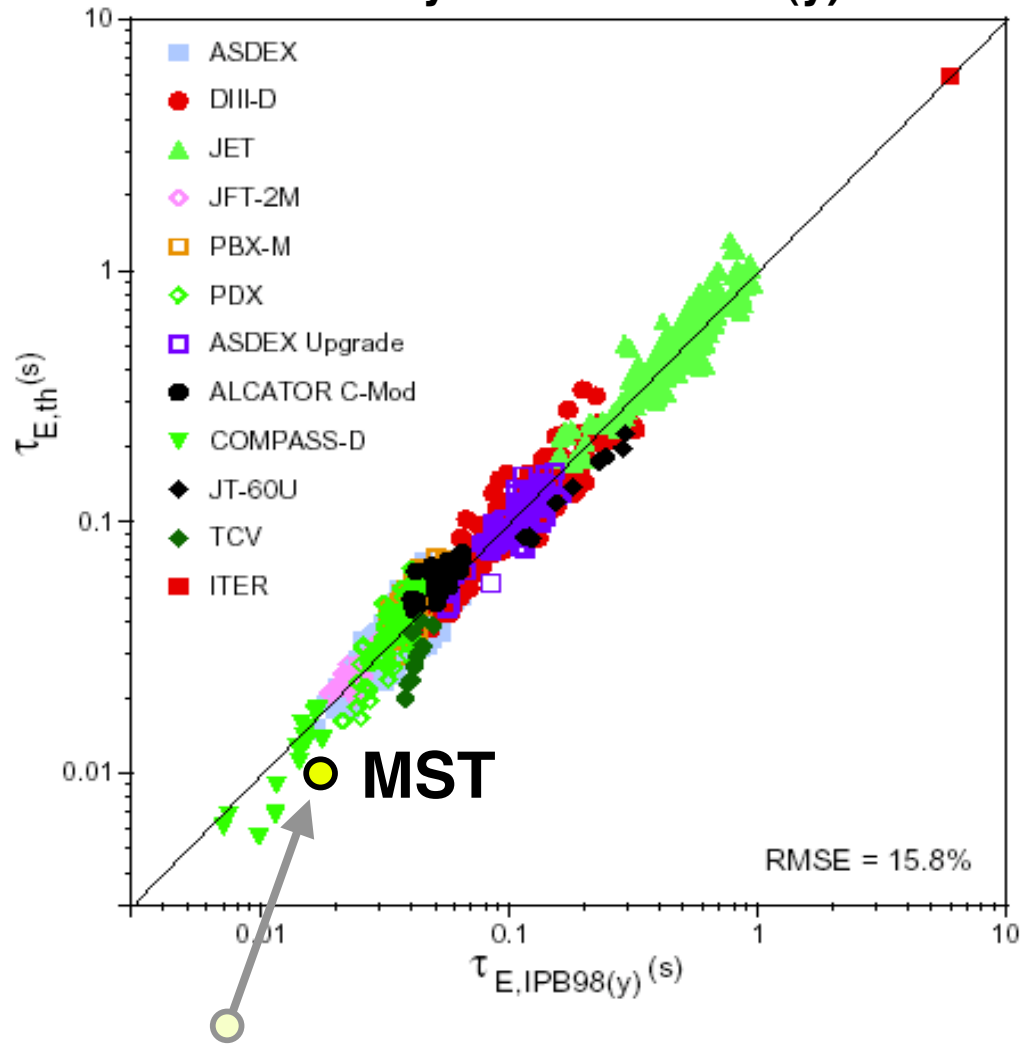
FIR  
interferometry

relate to tokamak database

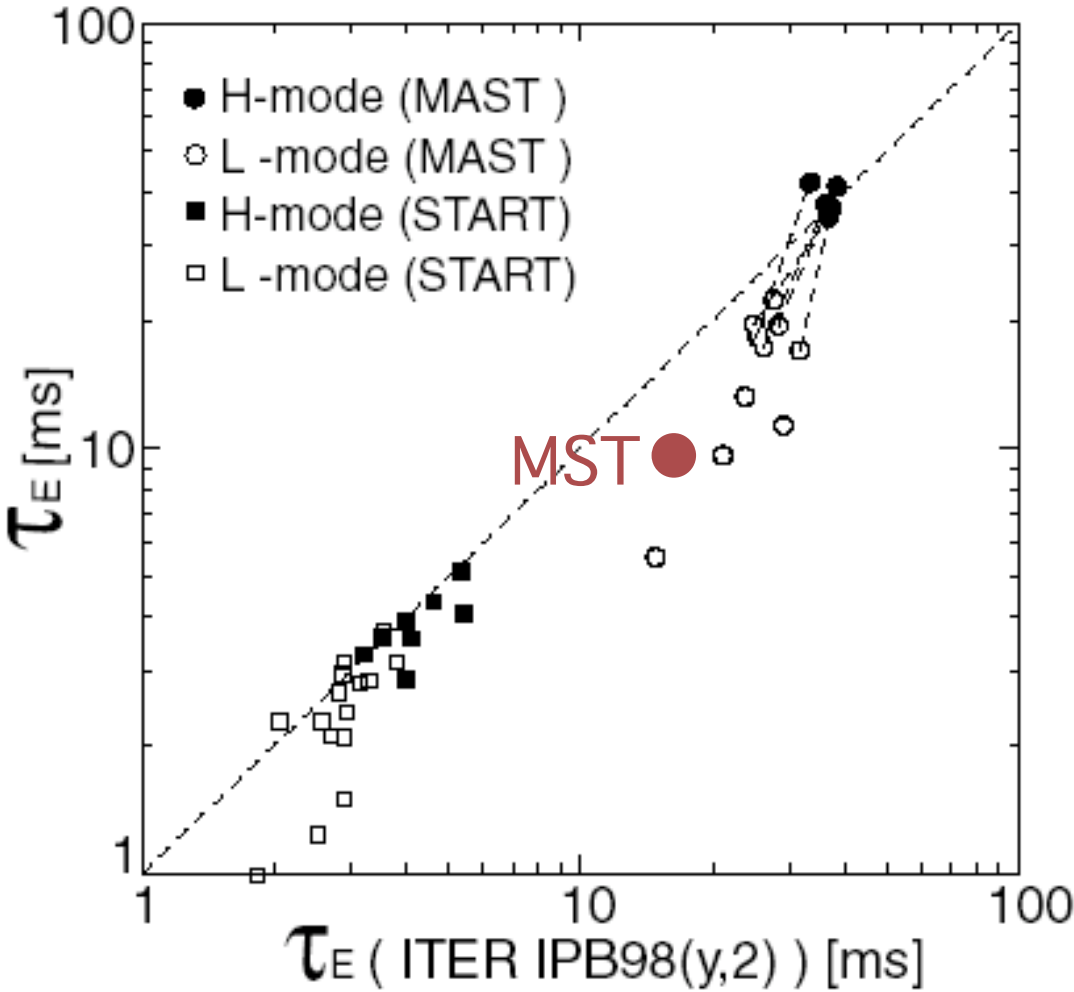




# ELMy H-Mode IPB89(y)

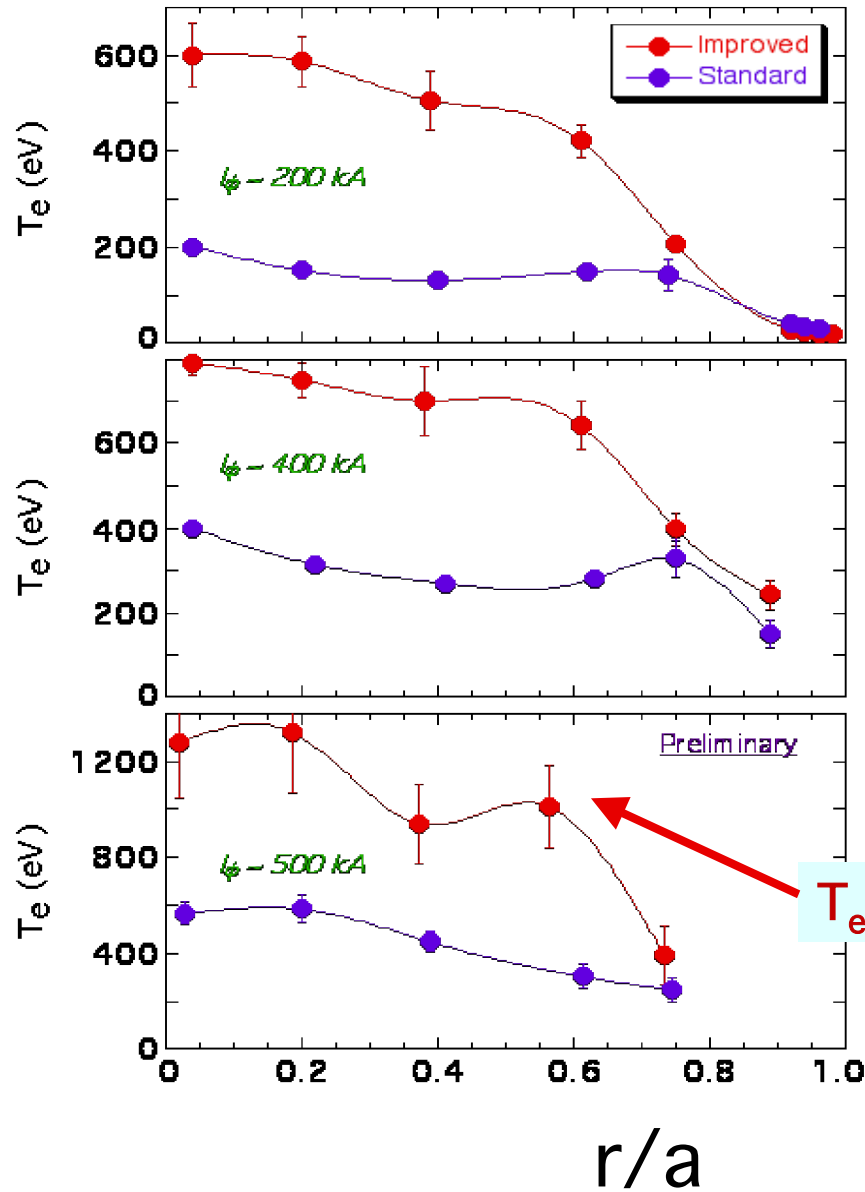


relate to current spherical tokamaks,



from MAST PRL

$T_e$  increases everywhere



$I = 200$  kA

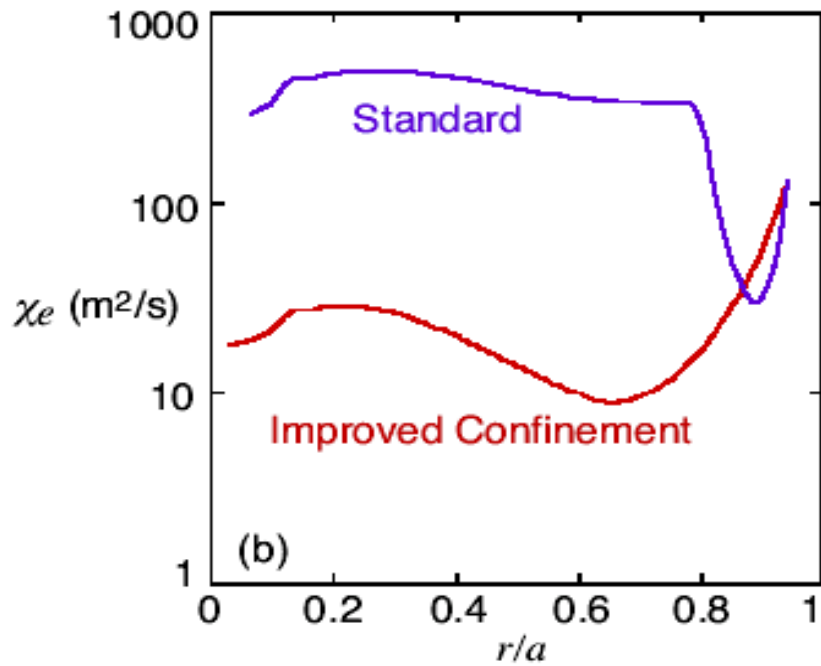
$I = 400$  kA

$I = 500$  kA

$T_e \rightarrow 1.3$  keV

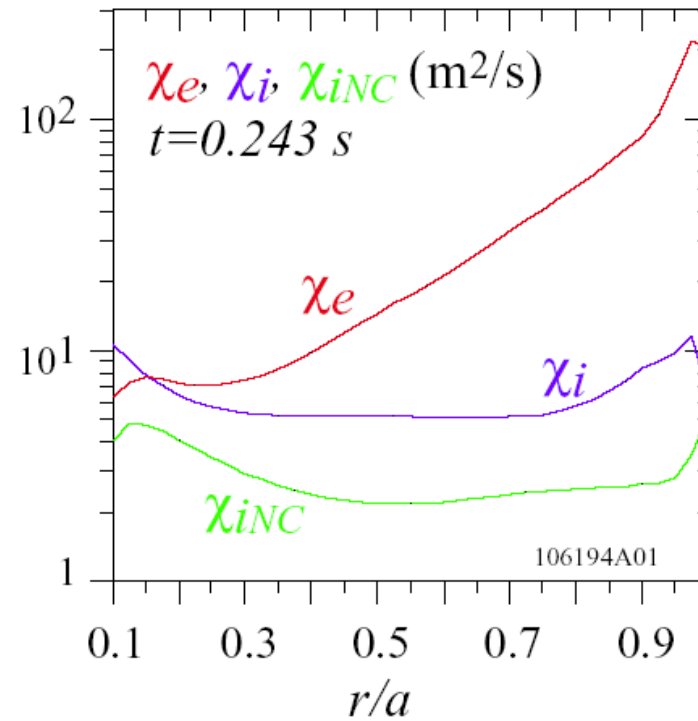
# MST electron thermal diffusivity similar to ST

## MST



*not best case (by 2x)*

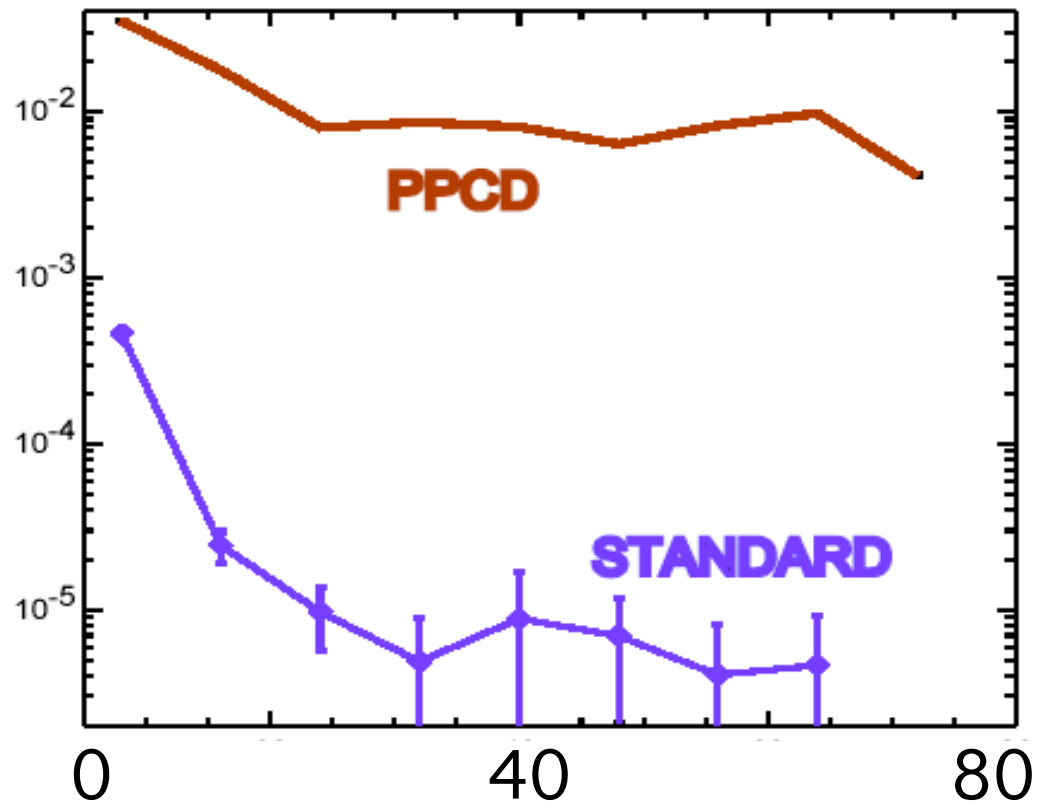
## NSTX



*From Wilson, ICC conference,  
likely not best case*

# Electrons confined to 100 keV

HXR  
flux



energy (keV)

Fokker-Planck modeling (CQL3D) implies

- Runaway electron density  $\sim 2\%$
- Electron travel distance  $\sim 100$  km
- Diffusion coefficient  $\sim 5$  m<sup>2</sup>/s
- **Magnetic surfaces are likely well-formed**

## MST beta values are increasing

$$\beta = \frac{\langle p \rangle_V}{\langle B^2 / 2\mu_0 \rangle_S} \sim 15\% \text{ in MST} \quad (\text{probably transport limited})$$

The “engineering beta”  $\frac{\langle p \rangle_V}{\langle B^2 / 2\mu_0 \rangle_{coil}}$  is higher than 15%

# MST Confinement Summary

(at  $I = 200$  kA)

|             | <u>standard</u>      |        | <u>PPCD</u>                 |
|-------------|----------------------|--------|-----------------------------|
| $T_e$       | 200 eV               | -----> | 600 eV<br>(1.3 keV @ 500kA) |
| Beta        | 9%                   | -----> | 15%                         |
| Ohmic power | 2 MW                 | -----> | 1 MW                        |
| $\tau_E$    | 1 ms                 | -----> | 10 ms                       |
| $\chi_e$    | 50 m <sup>2</sup> /s | -----> | 5 m <sup>2</sup> /s         |

*magnetic field no longer stochastic*



But,

- The **ultimate limit** to transport in the RFP is not yet understood
- Confinement improvement is not **sustained**

(although the projected PoP parameters are already nearly obtained transiently)

- MST is presently *intermediate* between a CE and PoP experiment
- We must sustain and thereby validate the improved confinement
- We must discover the ultimate RFP confinement
- This requires the PoP program approved by FESAC

## For refined, sustained plasma control:

- RF (LH, EBW)
- Oscillating field current drive
- Neutral beam injection
- Pellet injection (ORNL)

for current drive and heating;  
all are being initiated in stages

MST as a PoP program in plasma physics

*Proposal to NSF for*  
A Center for Magnetic Self-Organization in  
Laboratory and Astrophysical Plasmas

- To study high temperature plasma physics issues common to lab and cosmos
- Links lab and astrophysical scientists
- Links experiment, theory, computation
- Links four experiments (MST, MRX, SSPX, SSX)
- 24 participants from 7 institutions

# topics

- Dynamo
- Reconnection
- Magnetic helicity conservation and transport
- Angular momentum transport
- Ion heating
- Magnetic chaos and transport

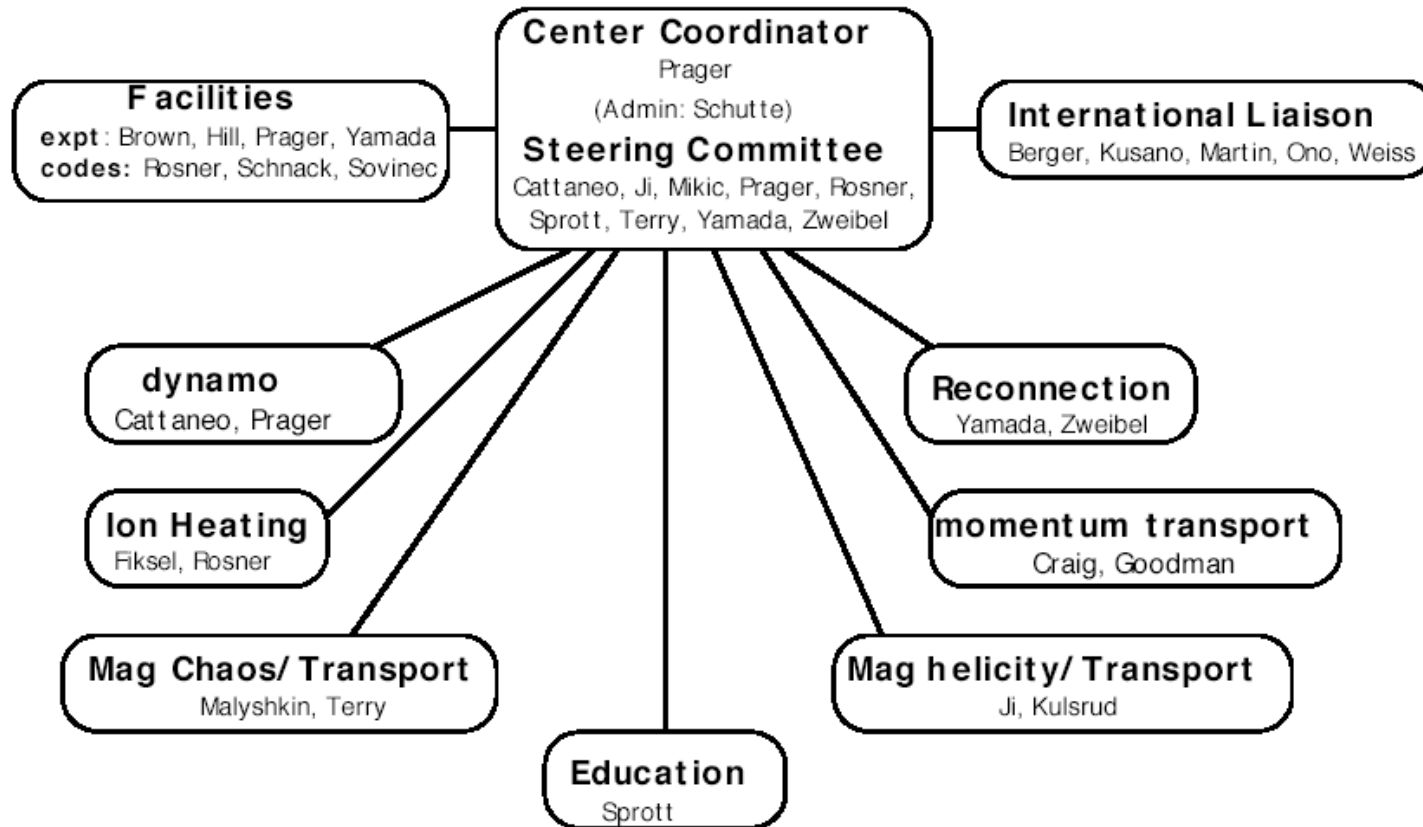
|  | <b>LABORATORY</b>  | <b>ASTROPHYSICS</b>   |
|--|--|---|
| <b>DYNAMO</b>                                      | RFP field sustainment,<br>Spheromak field sustainment,<br>RFP sawtooth crash/relaxation  | Solar magnetic field cycles, Earth magnetic field,<br>Stellar magnetic field cycles,<br>Accretion disk flux conversion  |
| <b>RECONNECTION</b>                                | Merging plasmas,<br>Spontaneous reconnection in RFP and spheromak,<br>Sawtooth oscillation,<br>Forced reconnection during helicity injection | Earth magnetosphere, Solar flares,<br>Star formation, Protostellar disks,<br>Particle acceleration to ultra-relativistic energy                                     |
| <b>HELICITY<br/>CONSERVATION<br/>AND TRANSPORT</b> | Relaxation/dynamo in RFP,<br>Relaxation/dynamo in spheromak,<br>Merging reconnection,<br>Helicity injection experiments                      | Disruptions in coronal loops,<br>Solar flares,<br>Helicity in solar wind,<br>Fast dynamo  |
| <b>ANGULAR<br/>MOMENTUM<br/>TRANSPORT</b>          | Momentum redistribution in the RFP,<br>Momentum generation in tokamaks   | Accretion disks of white dwarfs,<br>Accretion disks of black holes,<br>Accretion disks of AGN,<br>Differential rotation in the Sun,<br>Disks of non-accreting stars |
| <b>ION HEATING</b>                                 | RFP in steady-state, RFP during relaxation events,<br>Merging reconnection expts,<br>Spherical tokamak with neutral beam injection           | Solar corona and wind,<br>Earth magnetosphere,<br>Accretion flow onto black holes   |
| <b>MAGNETIC<br/>CHAOS AND<br/>TRANSPORT</b>        | Transport in RFP,<br>Transport in spheromak,<br>Transport during forced reconnection,<br>Kinetic dynamo in RFP, spheromak                    | Alfven waves in solar corona,<br>Heating in solar corona,<br>Cosmic ray transport in galactic magnetic field  |

# Participants

| Institution                                       | Participant   | Department                        | Expertise             |
|---|---------------|-----------------------------------|-----------------------|
| University of Chicago                             | F. Cattaneo   | Mathematics                       | Astro, comp           |
|   | T. Linde      | Astronomy & Astrophysics          | Astro, comp           |
|   | L. Malyshkin  | Astronomy & Astrophysics          | Astro, theory         |
|   | R. Rosner     | Astronomy & Astrophysics          | Astro, theory, comp   |
| Princeton University                              | J. Goodman    | Astrophysical Sciences            | Astro, theory         |
|   | H. Ji         | Plasma Physics Lab                | Lab, expt             |
|   | R. Kulsrud    | Astrophysical Sciences            | Lab & astro, theory   |
|   | M. Yamada     | Plasma Physics Lab                | Lab, expt             |
| University of Wisconsin<br>(The lead institution) | J. Cassinelli | Astronomy                         | Astro, theory, observ |
|   | D. Craig      | Physics                           | Lab, expt             |
|   | D. Den Hartog | Physics                           | Lab, expt             |
|   | G. Fiksel     | Physics                           | Lab, expt             |
|   | C. Hegna      | Engineering Physics               | Lab, theory           |
|   | A. Lazarian   | Astronomy                         | Astro, theory         |
|   | S. Prager     | Physics                           | Lab, expt             |
|   | J. Sarff      | Physics                           | Lab, expt             |
|   | C. Sovinec    | Engineering Physics               | Lab, comp             |
|   | C. Sprott     | Physics                           | Education, outreach   |
| P. Terry  | Physics       | Physics                           |                       |
| Individual Participants                           |               |                                   |                       |
| Swarthmore College                                | M. Brown      | Physics                           | Lab, expt             |
| Lawrence Livermore                                | D. Hill       |                                   | Lab, expt             |
| Science Applic Int Corp                           | Z. Mikic      |                                   | Astro, comp           |
| Science Applic Int Corp                           | D. Schnack    |                                   | Astro and lab, comp   |
| U. Colorado                                       | E. Zweibel    | JILA/ astrophys. & Planetary Sci. | Astro, theory         |



# Management Structure



# status

- Selected (from pre-proposal) to submit full proposal (12 out of 44 pre-proposals)
- Chance of NSF funding: about 25%?
- Validates idea that fusion research contributes to general science
- A PoP facility is needed for general plasma science (comprehensive diagnostics and physics)
- An opportunity to exploit

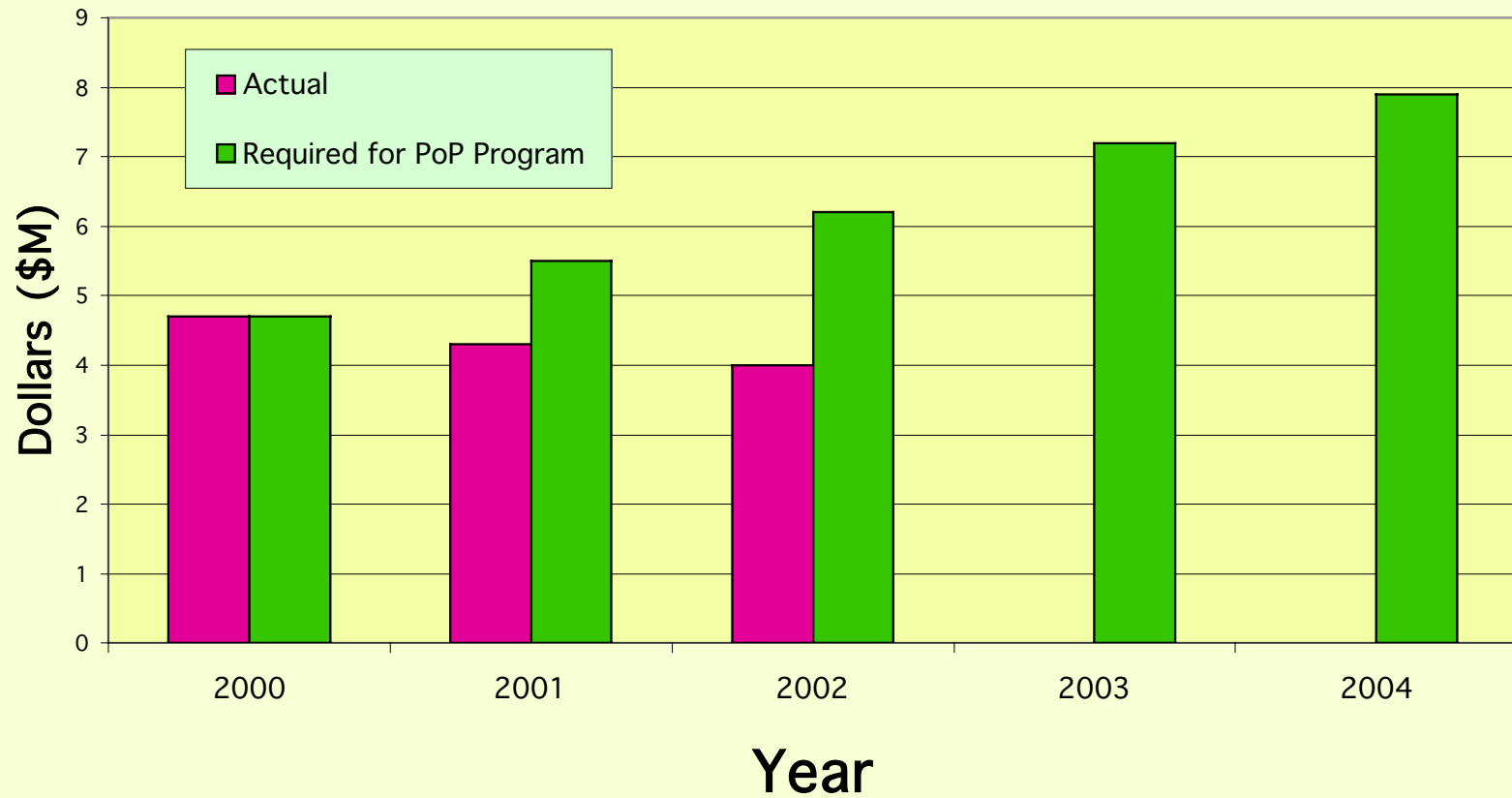
The RFP PoP program is easy to convey to policy makers

- An appealing scientific idea: control magnetic chaos
- Relevance to fusion: low field, high pressure
- Strong connection to broader science/astrophysics
- World leadership
- Extreme cost effectiveness (\$8M/yr)
- University participation at the PoP level
- Approved by rigorous review process

# MST Collaborators

- **RPI** - Heavy ion beam probe
- **UCLA** - FIR interferometry/polarimetry
- **Novosibirsk** - neutral beam diagnostics and heating
- **ORNL/GA** - lower hybrid/ EBW injection (help in planning expts)
- **ORNL** - pellet injection
- **University of Saskatchewan** - SXR, NBI expts
- **Aries group**- system studies
- **RFX, Italy** - Laser impurity injection, SXR, data analys
- **TPE-RX, Japan** -PPCD expts
- **Pegasus, HSX** - Thom scat, HXR detection, dnb

# The MST Funding Problem



# Summary

We are scientifically well-positioned to implement the 1999 FESAC recommendation to move forward with an RFP proof-of-principle program.

# MST Physics Group

**Wisconsin:** A. Almagri, J. Anderson, B. Chapman, P. Chattopadhyay, D. Craig, D. den Hartog, G. Fiksel, C. Forest, J. Goetz, K. McCollam, R. O'Connell, S. Prager, J. Reardon, J. Sarff, (*students: T. Biewer, A. Blair, S. Castillo, M. Cengher, S. Choi, D. Ennis, M. Wyman*)

**UCLA:** D. Brower, W. Ding, S. Terry

**RPI:** K. Connor, D. Demers, P. Schock

**Novosibirsk:** V. Davidenko, A. Ivanov